

# Worldsheet $\text{CFT}_2$ and Celestial $\text{CFT}_2$ : An $\text{AdS}_3/ \text{CFT}_2$ perspective

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# Introduction

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# Celestial conformal field theory

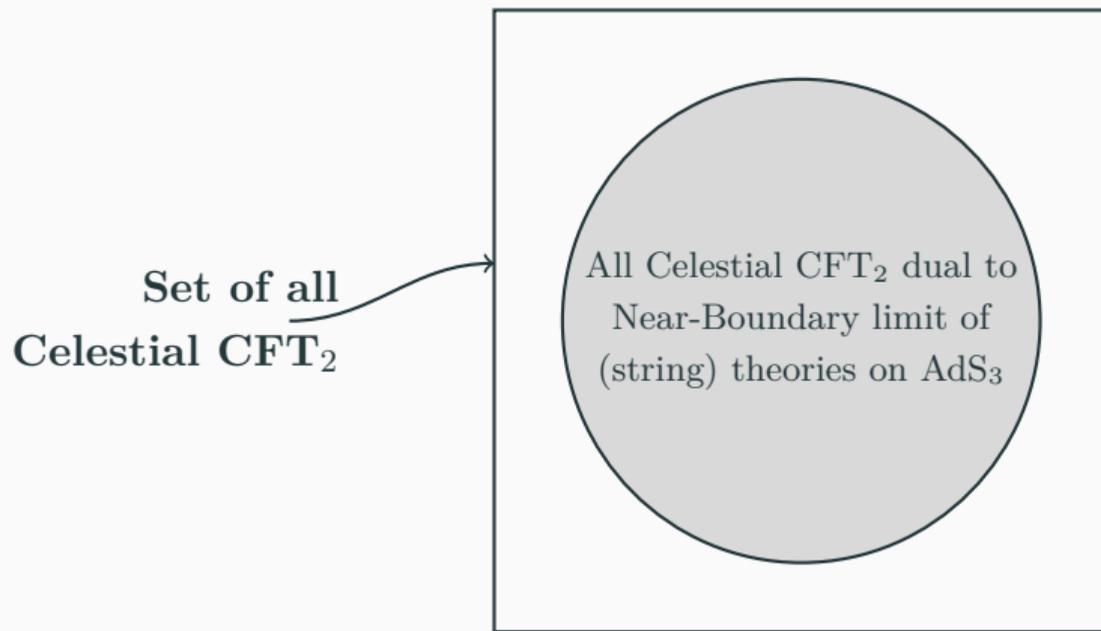
- Celestial holography: Quantum gravity in  $(d + 2)$  dimensional flat spacetime is dual to CFT on  $d$ -dimensional celestial sphere at the boundary null infinity.
- Symmetry group of Celestial  $\text{CFT}_d$  : Poincaré group  $\text{ISO}(d + 1, 1)$  includes conformal group  $\text{SO}(d + 1, 1)$  with translations realised as internal symmetry.
- Correlation functions compute the scattering amplitudes in  $(d + 2)$  flat space.
- Translations lead to distributional nature in correlation functions.
- Recent works consider  $\text{CCFT}_d$  with only Lorentz invariance ( $\text{SO}(d + 1, 1)$ ).  
[Melton, Sharma, Strominger]

# Celestial conformal field theory

- In these  $\text{CCFT}_d$ , correlation functions have the standard CFT form  $\dots$   
Reconstruct Poincaré invariant S-matrix elements from these correlation functions.
- Examples include Liouville theory coupled to conformal matter fields  $\dots$   
Celestial amplitudes for gluons and gravitons are holographically derived from correlation functions of Liouville vertex operators.  
[\[Melton-Sharma-Strominger-Wang, Stieberger-Taylor-Zhu, Donnay-Giribet-Valsesia\]](#)
- For this talk, Celestial  $\text{CFT}_d$  is any QFT with  $\text{ISO}(d+1, 1)$  invariance on which  $\text{SO}(d+1, 1)$  act as conformal group.

## In this talk

- A method to construct Celestial  $\text{CFT}_d$  from  $\text{AdS}_{d+1}/\text{CFT}_d$  duality  $\cdots$  **Scaling limit** where we zoom in the boundary of AdS.



- Introduction
- Zooming near the boundary of (EA)dS<sub>3</sub>
- Our Proposal
- Near Boundary limit of CFT two-point function
- Near boundary limit of massive field theory on EAdS<sub>3</sub>
- Near boundary limit of String theory
- Near boundary limit of Einstein Gravity
- Concluding Remarks

Zooming near the boundary of  
 $(EA)dS_3$

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# Conformal Killing Vector fields of (EA)dS<sub>3</sub>

- Consider CFT on (EA)dS<sub>3</sub> background

$$ds^2 = \frac{d\eta^2 + dz d\bar{z}}{\eta^2}, \quad \text{Boundary at } \eta = 0.$$

- Conformal Killing Vector fields  $\nabla_\mu \xi_\nu + \nabla_\nu \xi_\mu = \frac{2}{3} \nabla_\rho \xi^\rho g_{\mu\nu}$  10 parameters.

$$\mathcal{L}_m = -z^{m+1} \partial_z + \frac{1}{2} m(m+1) \eta^2 \partial_{\bar{z}} - \frac{1}{2} (m+1) z^m \eta \partial_\eta, \quad m \in \{0, \pm 1\}$$

$$\bar{\mathcal{L}}_m = -\bar{z}^{m+1} \partial_{\bar{z}} + \frac{1}{2} m(m+1) \eta^2 \partial_z - \frac{1}{2} (m+1) \bar{z}^m \eta \partial_\eta$$

$$\mathcal{P}_{r,s} = -z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \partial_\eta + 2\eta \left(s + \frac{1}{2}\right) z^{r+s} \partial_z + 2\eta \left(r + \frac{1}{2}\right) \bar{z}^{r+s} \partial_{\bar{z}} \\ + \left(r + \frac{1}{2}\right) \left(s + \frac{1}{2}\right) \eta^2 \partial_\eta, \quad (r, s) \in \left\{\pm \frac{1}{2}\right\}$$

# Conformal Killing Vector fields of (EA)dS<sub>3</sub>

- The vector fields obey  $\mathfrak{so}(4, 1)$  algebra

$$[\mathcal{L}_m, \mathcal{L}_n] = (m - n)\mathcal{L}_{m+n}, \quad [\bar{\mathcal{L}}_m, \bar{\mathcal{L}}_n] = (m - n)\bar{\mathcal{L}}_{m+n}$$

$$[\mathcal{L}_m, \mathcal{P}_{r,s}] = \frac{1}{2}(m - 2r)\mathcal{P}_{m+r,s}, \quad [\bar{\mathcal{L}}_m, \mathcal{P}_{r,s}] = \frac{1}{2}(m - 2s)\mathcal{P}_{r,m+s}$$

$$[\mathcal{P}_{r,s}, \mathcal{P}_{r',s'}] = 2(\epsilon_{rr'}\bar{\mathcal{L}}_{s+s'} + \epsilon_{ss'}\mathcal{L}_{r+r'}).$$

- Zooming in near the boundary  $\eta = \epsilon \tilde{\eta}$ ,  $\epsilon \rightarrow 0$  and  $\tilde{\eta}$  fixed
- In terms of  $\tilde{\eta}$  the CKVs become

$$\mathcal{L}_m = -z^{m+1} \partial_z - \frac{\tilde{\eta}}{2}(m+1)z^m \partial_{\tilde{\eta}} + \mathcal{O}(\epsilon^2)$$

$$\bar{\mathcal{L}}_m = -\bar{z}^{m+1} \partial_{\bar{z}} - \frac{\tilde{\eta}}{2}(m+1)\bar{z}^m \partial_{\tilde{\eta}} + \mathcal{O}(\epsilon^2)$$

$$\mathcal{P}_{r,s} = -z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \frac{\partial_{\tilde{\eta}}}{\epsilon} + \mathcal{O}(\epsilon).$$

## Group contraction of $\text{SO}(4, 1)$ to $\text{ISO}(3, 1)$

- In the limit  $\epsilon \rightarrow 0$ ,  $\mathcal{P}_{r,s}$  diverges  $\dots$  Define  $P_{r,s} = \epsilon \mathcal{P}_{r,s}$  and also denote

$$\mathcal{L}_m \xrightarrow{\epsilon \rightarrow 0} L_m, \quad \bar{\mathcal{L}}_m \xrightarrow{\epsilon \rightarrow 0} \bar{L}_m$$

- Resultant algebra becomes Poincare ( $\mathfrak{iso}(1, 3)$ ),

$$\begin{aligned} [L_m, L_n] &= (m - n)L_{m+n}, & [\bar{L}_m, \bar{L}_n] &= (m - n)\bar{L}_{m+n} \\ [L_m, P_{r,s}] &= \frac{1}{2} (m - 2r) P_{m+r,s}, & [\bar{L}_m, P_{r,s}] &= \frac{1}{2} (m - 2s) P_{r,m+s} \\ [P_{r,s}, P_{r',s'}] &= 0. \end{aligned}$$

- Near Boundary Limit:  $\text{SO}(4, 1)$  contracts to  $\text{ISO}(3, 1)$  with  $P_{r,s}$  as translations  $\dots$  Subgroup  $\text{SO}(3, 1)$  (isometry of  $\text{EAdS}_3$ ) remains unchanged.

## Symmetries of manifold obtained after near boundary limit

- Lorentz group  $SO(3, 1)$  acts on the coordinates  $(\tilde{\eta}, z, \bar{z})$  as

$$z \rightarrow \frac{az + b}{cz + d}, \quad \bar{z} \rightarrow \frac{\bar{a}\bar{z} + \bar{b}}{\bar{c}\bar{z} + \bar{d}}, \quad \tilde{\eta} \rightarrow \frac{\tilde{\eta}}{|cz + d|^2}.$$

and is the isometry group of the degenerate metric that we get after the near

boundary limit  $\boxed{ds_{\text{NB}}^2 = \frac{dzd\bar{z}}{\tilde{\eta}^2}}$

- The translation  $P_{r,s}$  acts as

$$z \rightarrow z, \quad \bar{z} \rightarrow \bar{z}, \quad \tilde{\eta} \rightarrow \tilde{\eta} + a_{r,s} z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}}, \quad (r, s) \in \{\pm \frac{1}{2}\}.$$

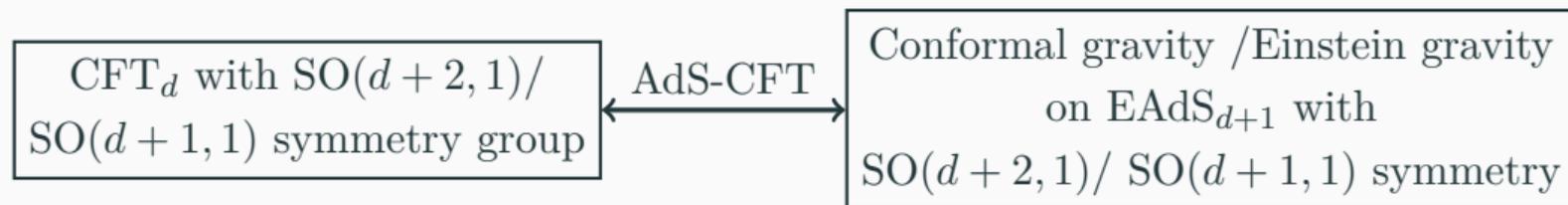
and act as (Weyl) conformal transformations.

# Our Proposal

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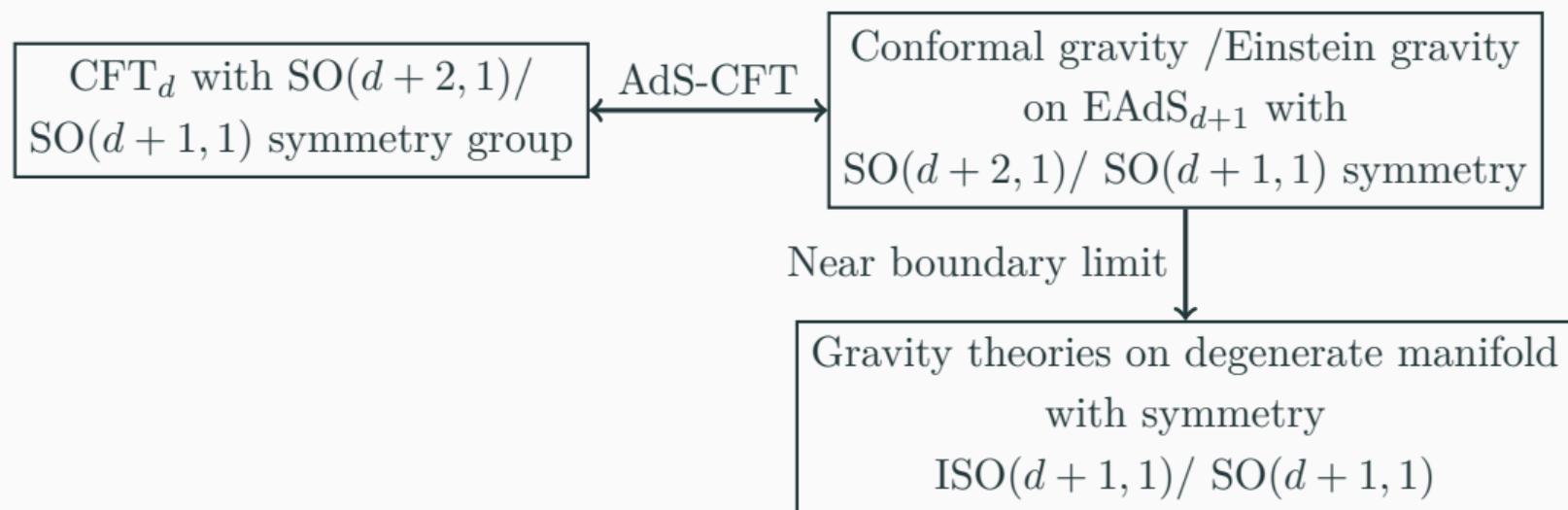
## Diagrammatic representation of the proposal

- In near boundary limit  $SO(d+2, 1)$  contracts to  $ISO(d+1, 1) \cdots$  Subgroup  $SO(d+1, 1)$  remain unchanged.



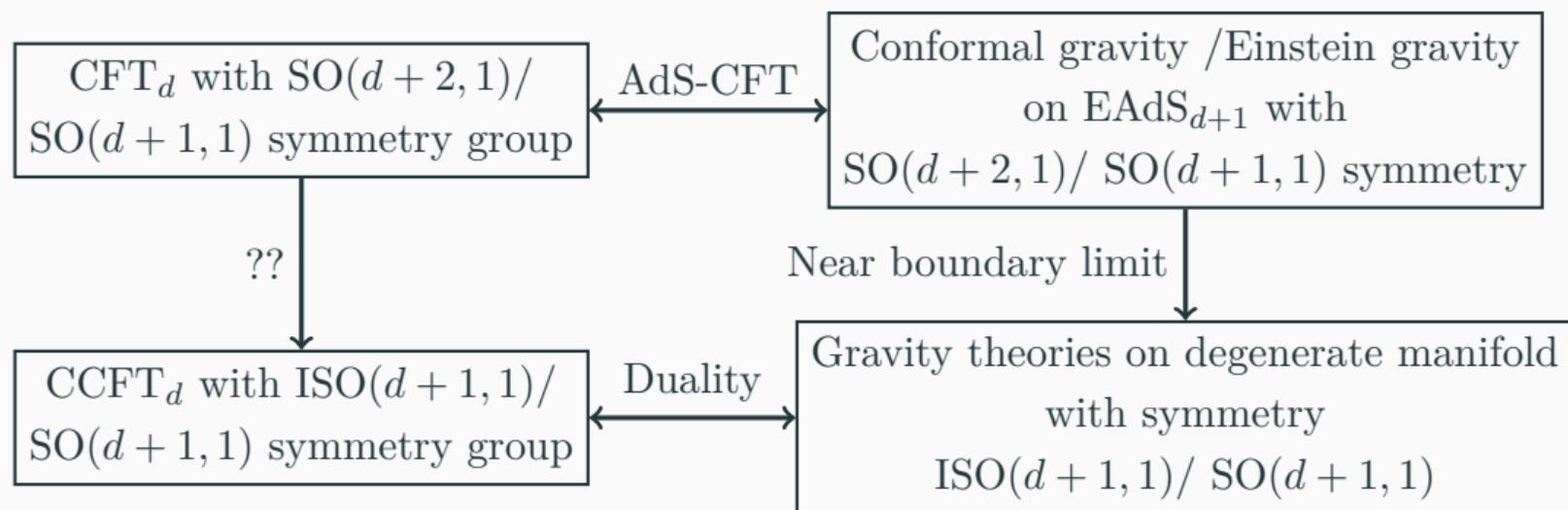
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- Argument extends to conformal and Einstein gravity on Lorentzian  $\text{AdS}_{d+1}$   
... Celestial  $\text{CFT}_d$  has symmetry group  $\text{ISO}(d, 2)$  and  $\text{SO}(d, 2)$ .
- Scaling limit does not change the bulk isometry algebra  $\text{SO}(d + 1, 1)$  ... Acts non-trivially on the correlation functions of bulk fields and therefore the bulk theory changes.
- Identification is based on only symmetries. Correlation functions of  $\text{CCFT}_d$  obtained this way may not necessarily compute scattering amplitudes in asymptotically flat  $(d + 2)$  dimensional spacetime.

# Near Boundary limit of CFT two-point function

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## Near Boundary limit of CFT two-point function

- Two-point function on  $EAdS_{d+1}$  conformally related to two-point function on flat space

$$\langle \Phi_{\Delta}(\eta_1, z_1, \bar{z}_1) \Phi_{\Delta}(\eta_2, z_2, \bar{z}_2) \rangle_{AdS_{d+1}} = \frac{\eta_1^{\Delta} \eta_2^{\Delta}}{\left( (\eta_{12})^2 + \vec{x}_{12}^2 \right)^{\Delta}}, \quad \vec{x} \in \mathbb{R}^d$$

- In the near boundary limit

$$\langle \Phi_{\Delta}(\epsilon \tilde{\eta}_1, \vec{x}_1) \Phi_{\Delta}(\epsilon \tilde{\eta}_2, \vec{x}_2) \rangle_{AdS_{d+1}} = \frac{\epsilon^{2\Delta} \tilde{\eta}_1^{\Delta} \tilde{\eta}_2^{\Delta}}{(\epsilon^2 \tilde{\eta}_{12}^2 + \vec{x}_{12}^2)^{\Delta}}$$
$$\rightarrow \epsilon^d \left( \pi^{\frac{d}{2}} \frac{\Gamma\left(\Delta - \frac{d}{2}\right)}{\Gamma(\Delta)} \frac{\tilde{\eta}_1^{\Delta} \tilde{\eta}_2^{\Delta}}{|\tilde{\eta}_{12}|^{2\Delta-d}} \delta^d(\vec{x}_1 - \vec{x}_2) \right) + \epsilon^{2\Delta} \left( \frac{\tilde{\eta}_1^{\Delta} \tilde{\eta}_2^{\Delta}}{|\vec{x}_1 - \vec{x}_2|^{2\Delta}} \right)$$

- Electric sector:**  $\Delta > \frac{d}{2}$ , **Magnetic sector:**  $\frac{d-1}{2} \leq \Delta \leq \frac{d}{2}$

- Defining rescaled field  $\Phi_{\Delta}(\epsilon \tilde{\eta}, \vec{x}) = \epsilon^{\frac{d}{2}} \tilde{\eta}^{\Delta} \tilde{\Phi}_{\Delta}(\tilde{\eta}, \vec{x})$
- In terms of  $\tilde{\Phi}$ , the two point function becomes

$$\langle \tilde{\Phi}_{\Delta}(\tilde{\eta}_1, \vec{x}_1) \tilde{\Phi}_{\Delta}(\tilde{\eta}_2, \vec{x}_2) \rangle_{\text{NB}} = \pi^{\frac{d}{2}} \frac{\Gamma\left(\Delta - \frac{d}{2}\right)}{\Gamma(\Delta)} \frac{\delta^d(\vec{x}_1 - \vec{x}_2)}{|\tilde{\eta}_{12}|^{2\Delta-d}}, \quad \Delta > \frac{d}{2}$$

- Two-point function of  $(d+1)$  Carrollian conformal primaries  $\tilde{\Phi}_{\Delta}(\tilde{\eta}, \vec{x})$  living on the degenerate metric

$$ds^2 = 0. d\tilde{\eta}^2 + d\vec{x}^2, \quad \vec{x} \in \mathbb{R}^d$$

- Not surprising since  $\text{ISO}(d+1, 1)$  is isomorphic to the conformal Carroll group in  $(d+1)$  dimensions.

## $\Delta > \frac{d}{2}$ as unitary bound in the electric sector

- As  $\Delta \rightarrow \frac{d}{2}+$ , finite part of the correlation function can be written as

$$\langle \tilde{\Phi}_\Delta(\tilde{\eta}_1, \vec{x}_1) \tilde{\Phi}_\Delta(\tilde{\eta}_2, \vec{x}_2) \rangle = -\frac{\pi^{\frac{d}{2}}}{\Gamma(\frac{d}{2})} \ln |\tilde{\eta}_1 - \tilde{\eta}_2|^2 \delta^d(\vec{x}_1 - \vec{x}_2).$$

- In electric sector of a Carrollian  $\text{CFT}_{d+1}$ , the scalar operator with dimension  $\Delta = \frac{d}{2}$  is not a conformal primary  $\dots$  derivative of scalar field is Carrollian conformal primary  $\dots$  similar to free scalar field with dimension zero.
- In a unitary Carrollian  $\text{CFT}_{d+1}$  the dimension  $\Delta$  of a scalar primary in the electric sector is bounded from below by  $\frac{d}{2}$ , i.e,  $\Delta > \frac{d}{2}$ .
- In a relativistic  $\text{CFT}_{d+1}$  the corresponding bound  $\Delta \geq \frac{d-1}{2}$  is less stringent.

## Magnetic sector

- For  $\Delta < \frac{d}{2}$ , the magnetic sector has a dominant contribution.
- In terms of the rescaled field  $\Phi_{\Delta}(\epsilon\tilde{\eta}, \vec{x}) = \epsilon^{\Delta} \tilde{\eta}^{\Delta} \tilde{\Phi}_{\Delta}(\tilde{\eta}, \vec{x})$  the two-point function becomes

$$\langle \tilde{\Phi}_{\Delta}(\tilde{\eta}_1, \vec{x}_1) \tilde{\Phi}_{\Delta}(\tilde{\eta}_2, \vec{x}_2) \rangle_{\text{NB}} = \frac{1}{|\vec{x}_1 - \vec{x}_2|^{2\Delta}}, \quad \frac{d-1}{2} \leq \Delta \leq \frac{d}{2}$$

- The lower bound on  $\Delta$  is the unitarity bound of the relativistic CFT $_{d+1}$ .  $\Delta$  of a scalar primary in the magnetic sector is both bounded from below and above by  $\frac{d-1}{2} \leq \Delta \leq \frac{d}{2}$ .
- In the near-boundary limit, conformal primary becomes electric Carrollian primary if  $\Delta > \frac{d}{2}$  and magnetic Carrollian primary if  $\frac{d-1}{2} \leq \Delta \leq \frac{d}{2}$ .

# Near boundary limit of massive field theory on EAdS<sub>3</sub>

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## Massive KG field action

- Consider scalar field of mass  $m$  on EAdS<sub>3</sub> with the following line element

$$ds^2 = d\phi^2 + e^{2\phi} d\gamma d\bar{\gamma}$$

where  $-\infty < \phi < \infty$  and  $\gamma$  and  $\bar{\gamma}$  are complex conjugates  $\dots$  boundary is at  $\phi = \infty$ .

- Relation with the old coordinates  $\phi = -\ln \eta$ ,  $\gamma = z$ ,  $\bar{\gamma} = \bar{z}$
- The action of the scalar field is given by

$$S = \int d\phi d\gamma d\bar{\gamma} e^{2\phi} \left( (\partial_\phi \Psi)^2 + e^{-2\phi} \partial_i \Psi \partial_i \Psi + m^2 \Psi^2 \right)$$

- In terms of  $\phi$  the near boundary limit becomes  $\boxed{\phi = -\ln \epsilon + \tilde{\phi}, \epsilon \rightarrow 0}$  with  $\tilde{\phi}$  held fixed.

## Near boundary limit of massive KG action

- In the near boundary limit, the action becomes

$$S_{\text{NB}} = \frac{1}{2} \int d\tilde{\phi} d\gamma d\bar{\gamma} e^{2\tilde{\phi}} \left( (\partial_{\tilde{\phi}} \tilde{\Psi})^2 + m^2 \tilde{\Psi}^2 \right)$$

- The action is invariant under  $\text{SO}(3, 1)$  which acts as follows

$$\tilde{\Psi}'(\tilde{\phi}', \gamma', \bar{\gamma}') = \tilde{\Psi}(\tilde{\phi}, \gamma, \bar{\gamma})$$

$$\tilde{\phi}' = \tilde{\phi} + \ln |c\gamma + d|^2, \quad \gamma' = \frac{a\gamma + b}{c\gamma + d}, \quad \bar{\gamma}' = \frac{\bar{a}\bar{\gamma} + \bar{b}}{\bar{c}\bar{\gamma} + \bar{d}}, \quad \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in \text{SL}(2, \mathbb{C})$$

- These coordinate transformations are the isometry of  $\boxed{ds^2 = e^{2\tilde{\phi}} d\gamma d\bar{\gamma}}$  on which the scalar field  $\tilde{\Psi}(\tilde{\phi}, \gamma, \bar{\gamma})$  lives.

- **Infinite Symmetries:** The  $SO(3,1)$  symmetries of the metric and the action can be enhanced to infinite-dimensional local symmetry of arbitrary holomorphic reparameterization

$$\tilde{\Psi}'(\tilde{\phi}', \gamma', \bar{\gamma}') = \tilde{\Psi}(\tilde{\phi}, \gamma, \bar{\gamma})$$

$$\tilde{\phi}' = \tilde{\phi} - \frac{1}{2} \ln \left| \frac{d\gamma'}{d\gamma} \right|^2, \quad \gamma' = \gamma'(\gamma), \quad \bar{\gamma}' = \bar{\gamma}'(\bar{\gamma})$$

- **Boundary Correlation function:** The boundary two-point correlation function for  $S_{\text{NB}}$  is computed using the bulk Green's function defined as

$$\left( -e^{-2\tilde{\phi}} \frac{\partial}{\partial \tilde{\phi}} \left( e^{2\tilde{\phi}} \frac{\partial}{\partial \tilde{\phi}} \right) + m^2 \right) G(\tilde{\phi}, \gamma, \bar{\gamma}; \tilde{\phi}', \gamma', \bar{\gamma}') = \frac{\delta(\tilde{\phi} - \tilde{\phi}') \delta^2(\gamma - \gamma')}{e^{2\tilde{\phi}}}$$

## Green's function for $S_{\text{NB}}$

- The Green's function  $G$  is

$$G(\tilde{\phi}, \gamma, \bar{\gamma}; \tilde{\phi}', \gamma', \bar{\gamma}') \\ = \frac{e^{-\Delta_+ \tilde{\phi}} e^{-\Delta_- \tilde{\phi}'}}{2\sqrt{1+m^2}} \theta(\tilde{\phi} - \tilde{\phi}') \delta^2(\gamma - \gamma') + \frac{e^{-\Delta_- \tilde{\phi}} e^{-\Delta_+ \tilde{\phi}'}}{2\sqrt{1+m^2}} \theta(\tilde{\phi}' - \tilde{\phi}) \delta^2(\gamma - \gamma')$$

where  $\Delta_{\pm} = 1 \pm \sqrt{1+m^2}$

- In terms of coordinate  $\tilde{\eta} = e^{-\tilde{\phi}}$

$$G(\tilde{\eta}, \gamma, \bar{\gamma}; \tilde{\eta}', \gamma', \bar{\gamma}') \\ = \frac{\tilde{\eta}^{\Delta_+} \tilde{\eta}'^{\Delta_-}}{2\sqrt{1+m^2}} \theta(\tilde{\eta}' - \tilde{\eta}) \delta^2(\gamma - \gamma') + \frac{\tilde{\eta}^{\Delta_-} \tilde{\eta}'^{\Delta_+}}{2\sqrt{1+m^2}} \theta(\tilde{\eta} - \tilde{\eta}') \delta^2(\gamma - \gamma')$$

## Boundary Correlation functions

- Taking  $(\tilde{\eta}, \tilde{\eta}')$  to zero and using the extrapolate dictionary, the boundary correlators are

$$\begin{aligned}\langle O_{\Delta_+}(\gamma, \bar{\gamma}) O_{\Delta_-}(\gamma', \bar{\gamma}') \rangle &= \delta^2(\gamma - \gamma'), \quad \Delta_+ + \Delta_- = 2 \\ \langle O_{\Delta_+}(\gamma, \bar{\gamma}) O_{\Delta_+}(\gamma', \bar{\gamma}') \rangle &= \langle O_{\Delta_-}(\gamma, \bar{\gamma}) O_{\Delta_-}(\gamma', \bar{\gamma}') \rangle = 0\end{aligned}$$

- The boundary two-point function is a pure contact term because the theory is ultralocal in the boundary direction.
- In the near boundary limit in  $\text{AdS}_{d+1}$  a scalar field of mass  $m$  corresponds to two boundary operators of dimensions

$$\Delta_{\pm} = \frac{d}{2} \pm \sqrt{\frac{d^2}{4} + m^2}.$$

# Near boundary limit of String theory

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## Worksheet theory for string propagation on $(\text{EA})\text{dS}_3 \times \mathcal{N}$

- Conventional string theories do not have target space conformal invariance  $\dots$   
The low energy effective theory is Einstein gravity.
- In the near boundary limit, string theories on AdS should be dual to Celestial CFTs with only conformal (Lorentz invariance) and not Poincare invariance.
- Bosonic string theory on  $\text{EAdS}_3 \times \mathcal{N}$  in the presence of Neveu-Schwarz  $B_{\mu\nu}$  field with  $\text{EAdS}_3$  metric  $ds^2 = l^2 (d\phi^2 + e^{2\phi} d\gamma d\bar{\gamma})$

- For specific form of  $B_{\mu\nu}$ , the worldsheet Lagrangian on  $\text{AdS}_3$  is given by

$$S = \frac{2l^2}{l_s^2} \int d^2w \left( \partial\phi\bar{\partial}\phi + e^{2\phi}\bar{\partial}\gamma\partial\bar{\gamma} \right) + S_{\text{int}}$$

- Bosonic part of Lagrangian is given by WZW model for coset  $\text{SL}(2, \mathbb{C})/\text{SU}(2)$   
...  $S_{\text{int}}$  describes string propagation on compact internal space.

[Giveon-Kutasov-Seiberg, de Boer-Ooguri-Robins-Tannenhauser, Maldacena-Ooguri]

- The action can be written as

$$S = \int d^2w \left( \partial\phi\bar{\partial}\phi - \frac{2}{\alpha_+} \sqrt{g} R\phi + \beta\bar{\partial}\gamma + \bar{\beta}\partial\bar{\gamma} - \beta\bar{\beta}e^{-\frac{2}{\alpha_+}\phi} \right) + S_{\text{int}}$$

- $(\beta, \gamma)$  holomorphic fields with weights  $(1, 0)$  and  $(0, 0)$  and  $(\bar{\beta}, \bar{\gamma})$

corresponding anti-holomorphic fields. ...

$$\alpha_+^2 = 2k - 4 = 2\frac{l^2}{l_s^2} - 4$$

[Giveon, Kutasov, Sieberg]

- In the near boundary limit, neglecting the last term, one gets the free worldsheet action,

$$S_{\text{NB}} = \int d^2w \left( \partial\tilde{\phi}\bar{\partial}\tilde{\phi} - \frac{2}{\alpha_+} \sqrt{g} R\tilde{\phi} + \beta\bar{\partial}\gamma + \bar{\beta}\partial\bar{\gamma} \right) + S_{\text{int}}$$

- Equations of motion of  $\beta(\bar{\beta})$  are  $\bar{\partial}\gamma = 0, \partial\bar{\gamma} = 0$ .
- $\gamma(\bar{\gamma})$  is a holomorphic (anti-holomorphic) map from the worldsheet to the boundary sphere ( $S^2$ ) in the target space. **These wrapped strings are known as the “long-strings”.**
- $S_{\text{NB}}$  has an affine  $SL_2 \times SL_2$  symmetry at level  $k$  generated by worldsheet currents  $J^a(z)$  and  $\bar{J}^a(\bar{z})$  with the central charge  $c = \frac{3k}{k-2}$ .
- The symmetries of the worldsheet action can be lifted to the space-time and become the Virasoro algebra of the dual  $\text{CFT}_2$ . [Giveon, Kutasov, Sieberg]

## Properties of long strings

- This spacetime dual CFT describe long strings with central charge  $c_{\text{spacetime}} = 6kp$ .  $p$  is number of times worldsheet wraps the boundary sphere. [Giveon, Kutasov, Sieberg]
- Seiberg and Witten argued that the space-time  $\text{CFT}_2$  describing long strings contains Liouville field  $\tilde{\phi}$  and modes of the fields describing location of the string in  $\mathcal{N}$  with central charge  $c_{\text{int}}$ .
- The Liouville field  $\tilde{\phi}$  has improvement term  $Q$  (for  $p = 1$ ) and central charge  $c_L$  and modes in compact space has central charge  $c_{\text{int}}$ .
- In order to find the improvement term we use  $c_L + c_{\text{int}} = 6k$

- Using the expression  $c_L = 1 + 3Q^2$  and  $c_{\text{int}} = 26 - \frac{3k}{k-2}$  one obtain

$$c_L = 1 + 3Q^2 = 1 + 6\frac{(k-3)^2}{k-2} \implies Q = (k-3)\sqrt{\frac{2}{k-2}}.$$

- The Liouville field describes the radial ( $\tilde{\phi}$ ) fluctuations of the long-string worldsheet.
- The space-time  $\text{CFT}_2$  dual to the string theory described by the worldsheet action  $S_{\text{NB}}$  is an example of a Celestial  $\text{CFT}_2$  (with only Lorentz invariance).

# Near boundary limit of Einstein Gravity

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## Einstein gravity on locally (EA)dS<sub>3</sub>

- Consider the metric of locally EAdS<sub>3</sub> in Fefferman-Graham gauge with the following line element

$$ds^2 = d\phi^2 + e^{2\phi} h_{ij}(\phi, x^i) dx^i dx^j, \quad x^i = \{\gamma, \bar{\gamma}\}.$$

- On a constant  $\phi$  hypersurface, the two dimensional induced boundary metric

$$a_{ij} = e^{2\phi} h_{ij}.$$

- The Einstein-Hilbert action for EAdS<sub>3</sub> gravity becomes,

$$S_{\text{EH}} = \int d\phi d^2\gamma e^{2\phi} \sqrt{h} (K^2 - K_{ij}K^{ij} + 2) + \int d\phi d^2\gamma \sqrt{h} \mathcal{R}^{(2)}$$

- $\mathcal{R}^{(2)}$  is the Ricci scalar of metric  $h_{ij}$  and extrinsic curvature  $K_{ij}$  is extrinsic

curvature 
$$K_{ij} = \frac{1}{2} \partial_\phi a_{ij}, \quad K = a^{ij} K_{ij}.$$

## Near boundary limit of Einstein gravity

- In the near boundary limit, the action is

$$S_{\text{NB}} = \int d\tilde{\phi} d^2\gamma e^{2\tilde{\phi}} \sqrt{h} \left( \tilde{K}^2 - \tilde{K}_{ij} \tilde{K}^{ij} + 2 \right), \quad -\infty < \tilde{\phi} < \infty$$

with  $\tilde{K}_{ij} = \frac{1}{2} \partial_{\tilde{\phi}} \left( e^{2\tilde{\phi}} h_{ij}(\tilde{\phi}, x^i) \right)$ .

- The theory in the near boundary limit of Einstein gravity is defined on the degenerate manifold with line element

$$ds^2 = 0 \cdot d\tilde{\phi}^2 + e^{2\tilde{\phi}} h_{ij}(\tilde{\phi}, x^i) dx^i dx^j.$$

- $S_{\text{NB}}$  matches with the Carrollian limit of the Einstein-Hilbert action. [ [Hansen, Obers, Oling and Sjøgaard](#) ]

# Properties of near boundary limit of Einstein gravity

- The solution to equations of motion obtained from  $S_{\text{NB}}$  are exact

$$h_{ij}(\phi, x^i) = h_{ij}^{(0)}(x^i) + e^{-2\phi} h_{ij}^{(1)}(x^i) + e^{-4\phi} h_{ij}^{(2)}(x^i), \quad h_{ij}^{(2)} = \frac{1}{4} h_{ik}^{(1)} h_{(0)}^{kl} h_{lj}^{(1)}.$$

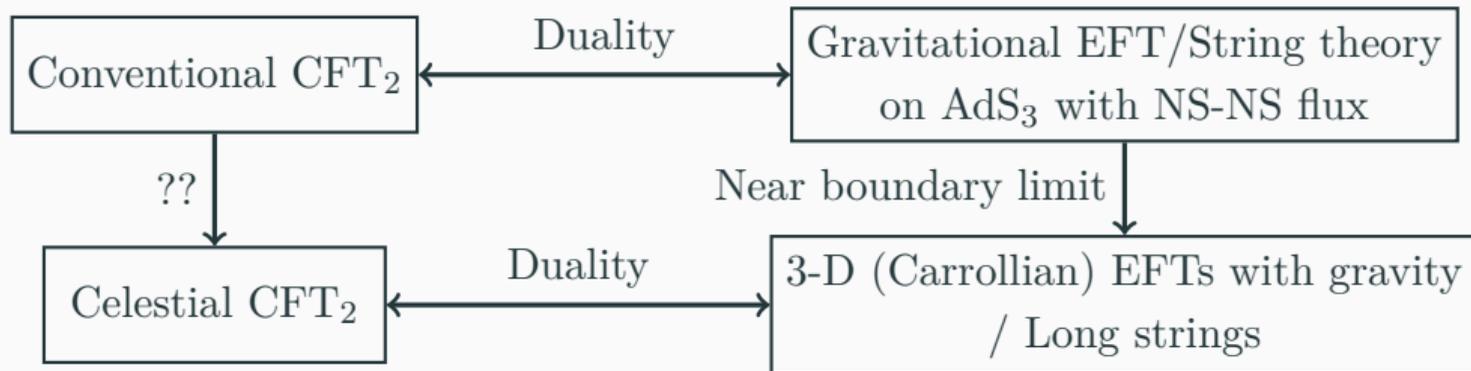
$$\boxed{h_{(0)}^{ij} h_{ij}^{(1)} = 0, \quad D^k h_{ki}^{(1)} = 0.}$$

- The tracelessness of  $h_{ij}^{(1)}$  implies zero central charge.
- Conservation of  $h_{ij}^{(1)}$  implies the boundary stress tensor leads to two commuting copies of Virasoro algebra.

## Concluding Remarks

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# Conclusions



- We focused on theories with Lorentz invariance  $SO(1, 3)$ . These results can be extended to conformal field theories and conformal gravity on  $(EA)dS_3$  (work in progress)
- String theory with target space conformal invariance?

- The existence of the Liouville sector in the long-string  $\text{CFT}_2$  suggests that the Liouville sector could be a universal feature of all Celestial  $\text{CFT}_2$ .
- Our observation resonates with the fact that in the high-energy limit in asymptotically flat space-time, one can identify the string worldsheet with the celestial sphere. [\[Stieberger, Taylor\]](#)

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Thank You

## Appendix A

- For the quantities defined as follows

$$\zeta_{ij} = \frac{1}{2}\partial_\phi h_{ij}, \quad \Theta = h^{ij}\zeta_{ij}$$

- The EE in the near boundary limit becomes

$$\Theta - \frac{1}{2}\zeta_{ij}\zeta^{ij} + \frac{1}{2}\Theta^2 = 0$$

$$h^{kl}\nabla_k\zeta_{il} - h^{kl}\nabla_i\zeta_{kl} = 0$$

$$\left(\partial_{\bar{\phi}}\Theta h_{ij} - \partial_{\bar{\phi}}\zeta_{ij} + \zeta_{kl}\zeta^{kl} h_{ij} + \zeta_i^k\zeta_{kj} + 2\Theta h_{ij} - 2\zeta_{ij}\right) = 0$$

All indices are raised and lowered by  $h_{ij}$ .

## Digression: Carrollian CFT primer

- Carrollian holography is another candidate of flat space holography.
- The dual to quantum gravity on  $(d + 2)$  dimensional asymptotically flat spacetime is  $(d + 1)$  dimensional Carrollian conformal field theory defined on boundary null infinity.
- Null infinity is an example of a Carrollian manifold, which is endowed with a degenerate metric and non-vanishing vector field. For example, in Bondi gauge, at null infinity, the boundary metric of  $(d + 2)$ -dimensional asymptotically flat spacetime is

$$ds^2 = 0 \cdot du^2 + dx^i dx^i, \quad \vec{x} \in \mathbb{R}^d$$

- The conformal group of Carrollian manifold is isomorphic to the Poincaré group  $\text{ISO}(d + 1, 1)$ .

## Transformation of $\Phi_\Delta$ under $\text{SO}(4, 1)$

- Transformation of 3d conformal scalar primary under  $\text{SO}(4, 1)$

$$\delta\Phi_\Delta = -\xi^\mu \partial_\mu \Phi_\Delta - \frac{\Delta}{3} \nabla_\mu \xi^\mu \Phi_\Delta$$

- The vector field  $\xi = a_m \mathcal{L}_m + \bar{a}_m \bar{\mathcal{L}}_m + p_{r,s} \mathcal{P}_{r,s}$
- Generator  $Q_\xi$  of the conformal symmetry  $[Q_\xi, \Phi_\Delta(\eta, z, \bar{z})] = \delta\Phi_\Delta$

## Transformation of $\Phi_\Delta$ under $\text{SO}(4, 1)$

- The commutation relation between the generators  $\text{SO}(4,1)$  and  $\Phi_\Delta(\eta, z, \bar{z})$

$$[\mathcal{L}_m, \Phi_\Delta] = z^{m+1} \partial_z \Phi_\Delta + \frac{1}{2}(m+1) z^m \eta \partial_\eta \Phi_\Delta - \frac{1}{2} m(m+1) \eta^2 \partial_{\bar{z}} \Phi_\Delta$$

$$[\bar{\mathcal{L}}_m, \Phi_\Delta] = \bar{z}^{m+1} \partial_{\bar{z}} \Phi_\Delta + \frac{1}{2}(m+1) \bar{z}^m \eta \partial_\eta \Phi_\Delta - \frac{1}{2} m(m+1) \eta^2 \partial_z \Phi_\Delta$$

$$[\mathcal{P}_{r,s}, \Phi_\Delta] = z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \left( \partial_\eta \Phi_\Delta - \frac{\Delta \Phi_\Delta}{\eta} \right) - \left( 2\eta \left( s + \frac{1}{2} \right) z^{r+s} \partial_z \Phi_\Delta \right. \\ \left. + 2\eta \left( r + \frac{1}{2} \right) \bar{z}^{r+s} \partial_{\bar{z}} \Phi_\Delta + \left( \left( r + \frac{1}{2} \right) \left( s + \frac{1}{2} \right) \right) (\eta^2 \partial_\eta \Phi_\Delta + \Delta \eta \Phi_\Delta) \right)$$

- $m \in \{0, \pm 1\}$  and  $(r, s) \in \{\pm \frac{1}{2}\}$
- Rescaling  $\eta = \epsilon \tilde{\eta}$  and defining Carrollian conformal primary field  $\tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z})$  as

$$\Phi_\Delta(\epsilon \tilde{\eta}, z, \bar{z}) = \tilde{\eta}^\Delta \epsilon \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z}), \quad \epsilon \rightarrow 0$$

## Transformation of $\tilde{\Phi}_\Delta$ under $\text{ISO}(3, 1)$

- In the limit  $\epsilon \rightarrow 0$ , the commutator becomes

$$[L_m, \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z})] = z^{m+1} \partial_z \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z}) + (m+1) z^m \left( \frac{\tilde{\eta}}{2} \partial_{\tilde{\eta}} + \frac{\Delta}{2} \right) \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z}) + \mathcal{O}(\epsilon^2)$$

$$[\bar{L}_m, \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z})] = \bar{z}^{m+1} \partial_{\bar{z}} \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z}) + (m+1) \bar{z}^m \left( \frac{\tilde{\eta}}{2} \partial_{\tilde{\eta}} + \frac{\Delta}{2} \right) \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z}) + \mathcal{O}(\epsilon^2)$$

$$[P_{r,s}, \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z})] = z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \partial_{\tilde{\eta}} \tilde{\Phi}_\Delta(\tilde{\eta}, z, \bar{z}) + \mathcal{O}(\epsilon^2)$$

- This is transformation law of a scalar primary field of dimension  $\Delta$  in a Carrollian  $\text{CFT}_3$  which lives on the manifold with degenerate metric

$$ds^2 = 0 \cdot d\tilde{\eta}^2 + dz d\bar{z}.$$